# An *In Situ* Approach for Approximating Complex Computer Simulations and Identifying Important Time Steps

# Kary Myers, Statistical Sciences Group Los Alamos National Laboratory

Joint work with Earl Lawrence, Mike Fugate, Claire Bowen, Larry Ticknor, Joanne Wendelberger, Jon Woodring, and Jim Ahrens

Sponsored by the National Nuclear Security Administration under contract DE-AC52-06NA25396.

LA-UR-15-23050.





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Join us for the third Conference on Data Analysis, bringing together scientists, statisticians, and data analysts from across the Department of Energy national laboratories along with their academic and industrial collaborators.

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- Subsurface Modeling
- Multisource Data

- Cyber Security
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- Really Expensive Data

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# Important announcement #2: Statistics and Beer Day

A new holiday to celebrate how the field of statistics has improved the world by focusing on how it has improved beer.

When? June 13, William Sealy Gosset's birthday.

**Who?** You might remember him by his pseudonym Student, as in Student's *t*-distribution.

**Why?** Gosset worked for Guinness Brewery where he developed and applied statistical methods to improve the beer.

**How?** Have a few pints with statisticians and other normal people. Start with a Guinness.



commons.wikimedia.org

Slide 3



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LA-UR-15-23050



# An *In Situ* Approach for Approximating Complex Computer Simulations and Identifying Important Time Steps!

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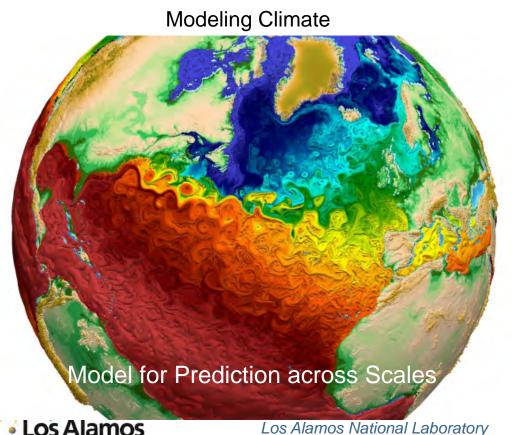
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LA-UR-15-23050.

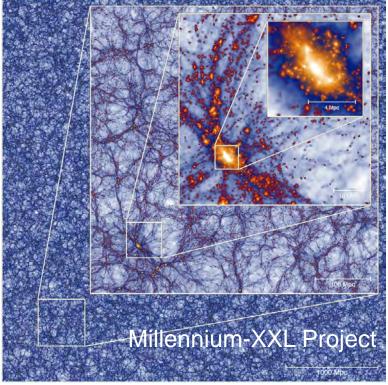


## What's a complex computer simulation?

Running a numerical model of a system, typically on a supercomputer. Often used when physical experiments aren't practical.



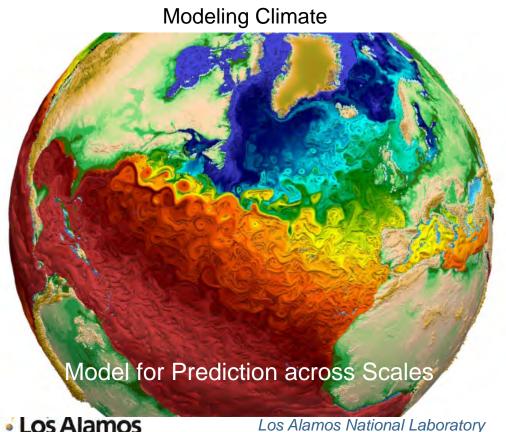
Modeling the Universe



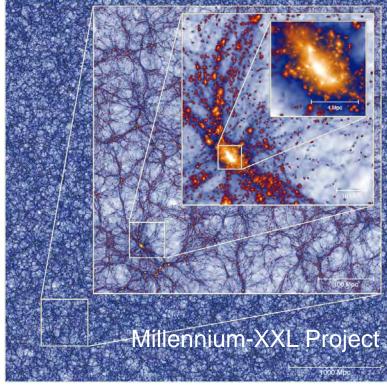
Max-Planck-Institute for Astrophysics

# **Coming soon: Exascale computing**

A billion billion (1,000,000,000,000,000) calculations every second. This means more science, but only if we can extract useful information.



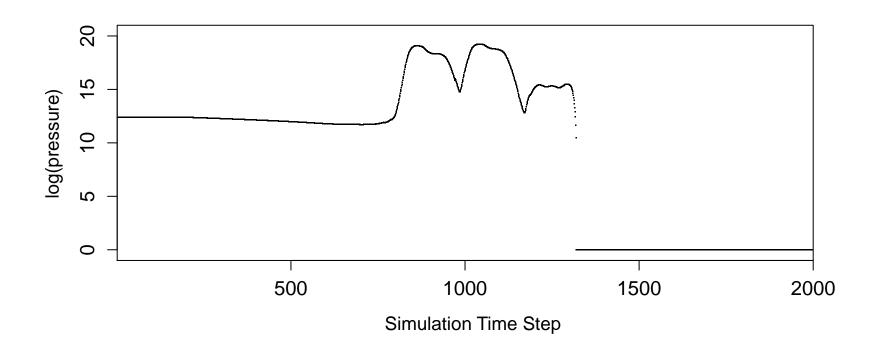
Modeling the Universe



Max-Planck-Institute for Astrophysics

# One idea: Only save a subset of simulation time steps

A 1-d example. (I'll talk about the simulation behind this example later.)

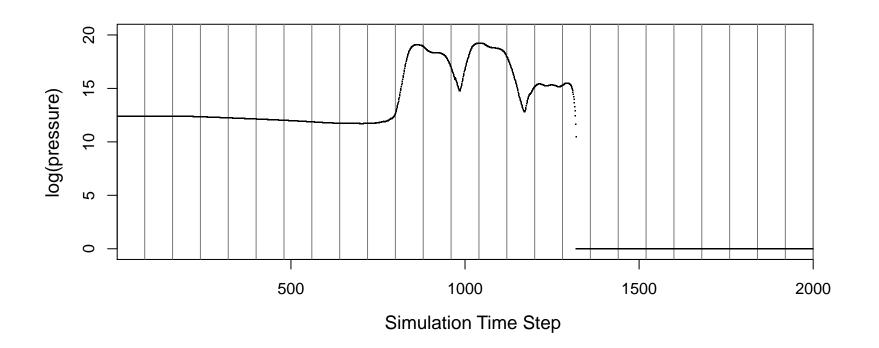






# One idea: Only save a subset of simulation time steps

Standard practice: Save evenly spaced time steps.

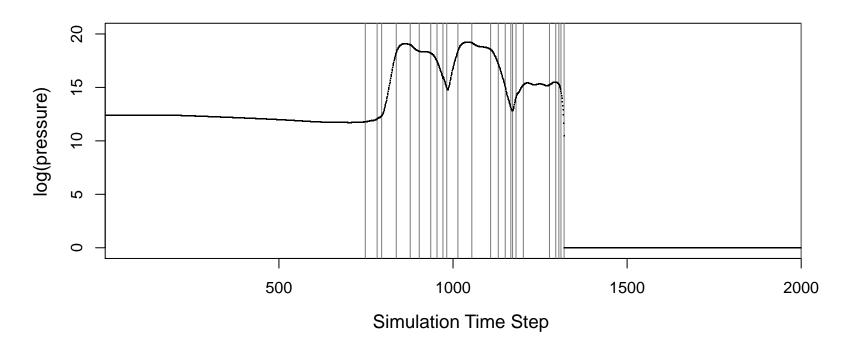






# One idea: Only save a subset of simulation time steps

Our approach: An *in situ* analysis to select "important" time steps in an online fashion. We do this by cheaply computing and comparing linear fits.

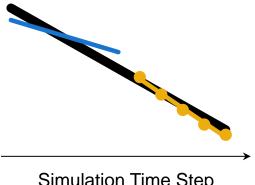




Cheap to compute and update within the simulation as it's running.

#### Compare 3 lines in 2 temporal regions of interest:

- curr: Time steps currently characterized by a linear fit; only sufficient statistics are stored.
- buff: B time steps most recently computed by the simulation; stored in the buffer.



Simulation Time Step

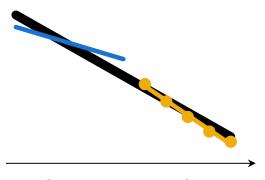


Cheap to compute and update within the simulation as it's running.

#### **Consider 2 hypotheses:**

 $H_0$ : **One line** fits best.

 $H_1$ : Two lines (curr + buff) fit best.



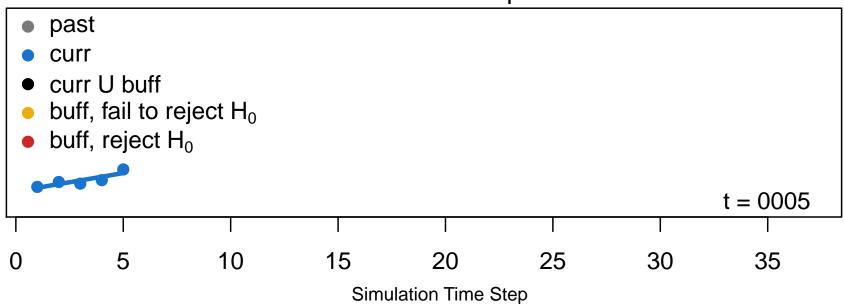
Use a **modified** *F*-statistic at some  $\alpha$  level to decide when to reject  $H_0$ .

Simulation Time Step



A toy example: Piecewise linear data with noise. Buffer size B = 5.

#### Fit a line to the time steps in curr

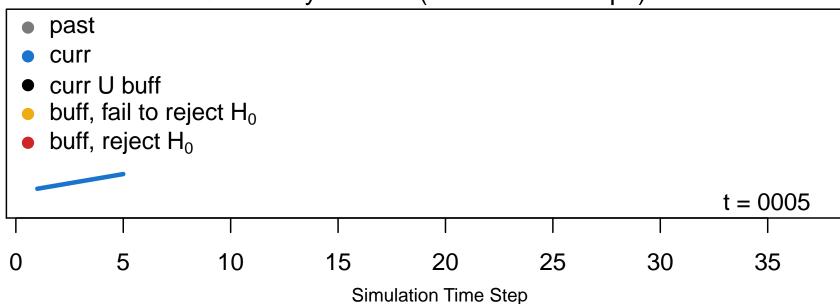






A toy example: Piecewise linear data with noise. Buffer size B = 5.

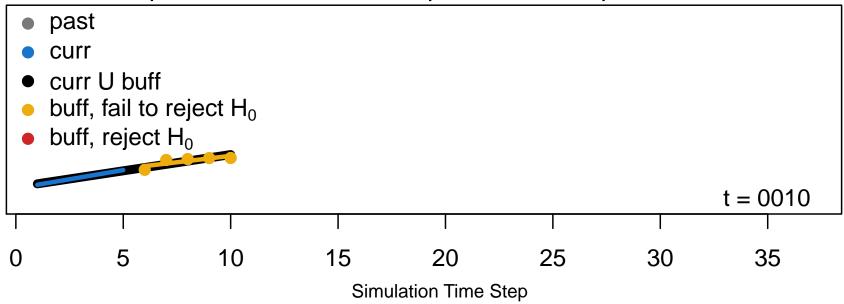
#### Save only the line (not the time steps)





A toy example: Piecewise linear data with noise. Buffer size B = 5.

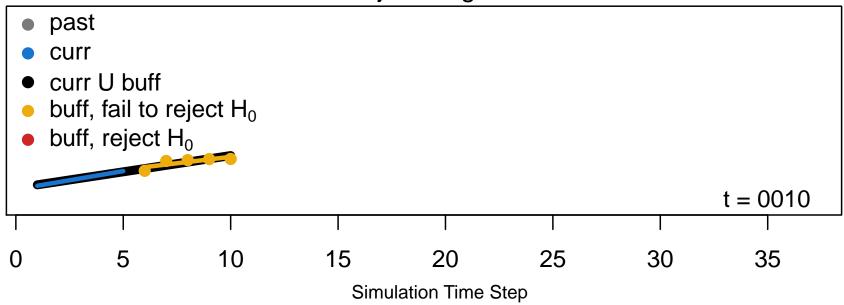
#### Acquire the next B time steps; fit and compare 3 lines







A toy example: Piecewise linear data with noise. Buffer size B = 5.

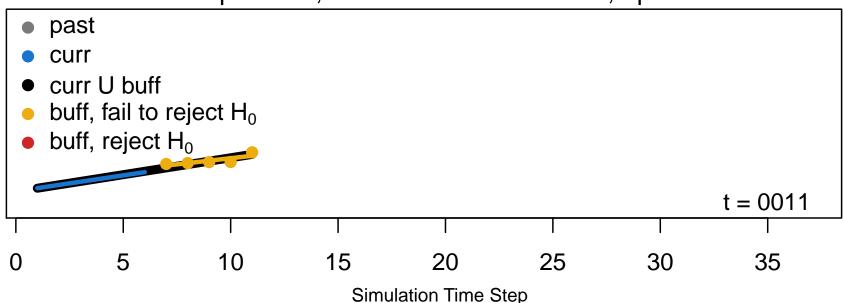




A toy example: Piecewise linear data with noise. Buffer size B = 5.

And throw it away!

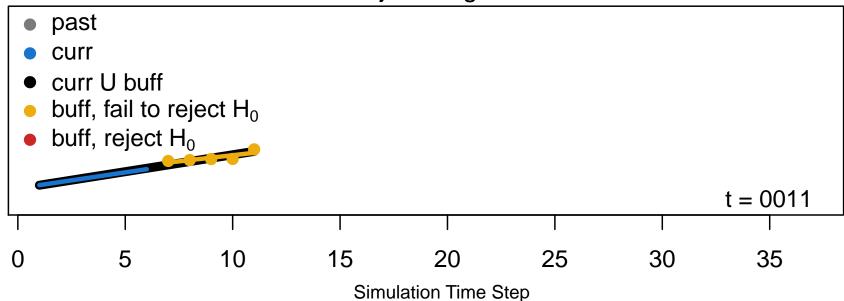
Add new time step to buff, move oldest one to curr, update the 3 lines





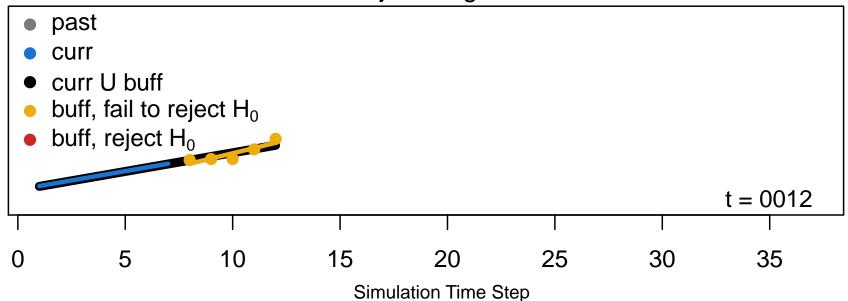


A toy example: Piecewise linear data with noise. Buffer size B = 5.





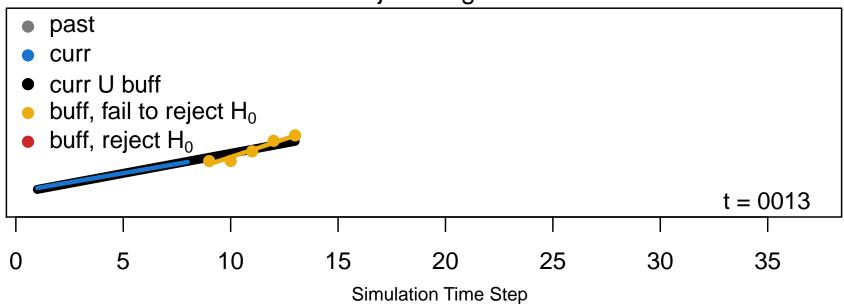
A toy example: Piecewise linear data with noise. Buffer size B = 5.







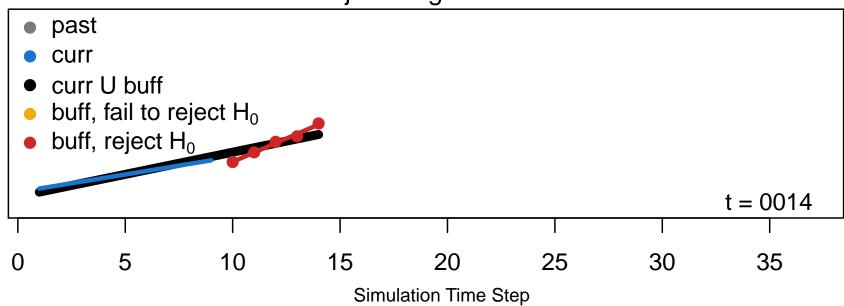
A toy example: Piecewise linear data with noise. Buffer size B = 5.



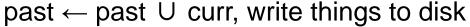


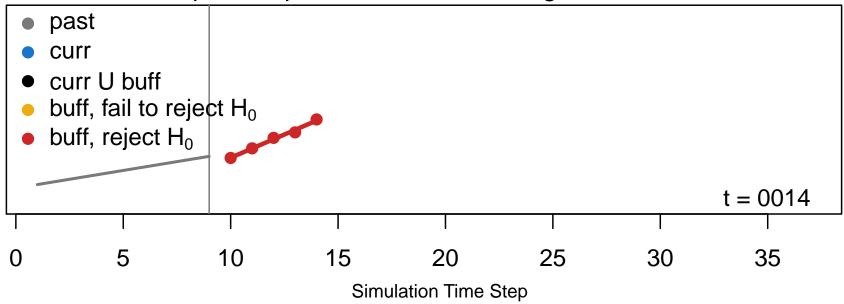
A toy example: Piecewise linear data with noise. Buffer size B = 5.

#### Reject single line fit

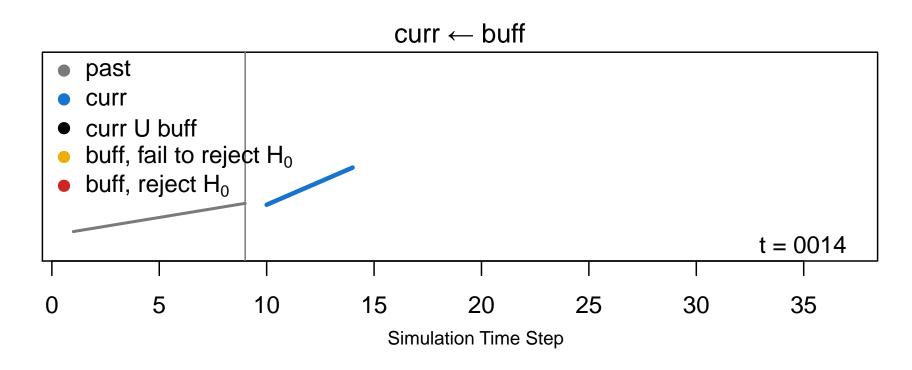








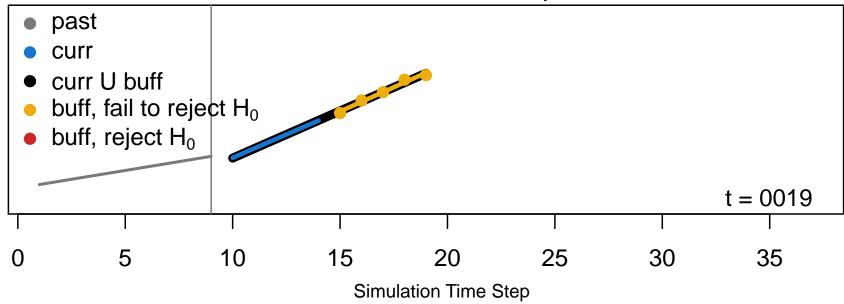




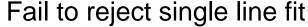


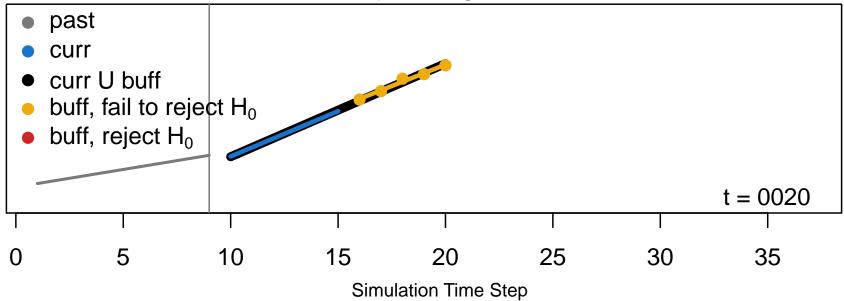






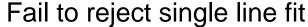


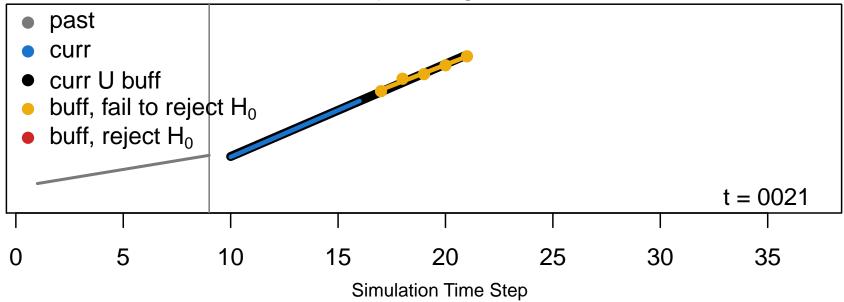






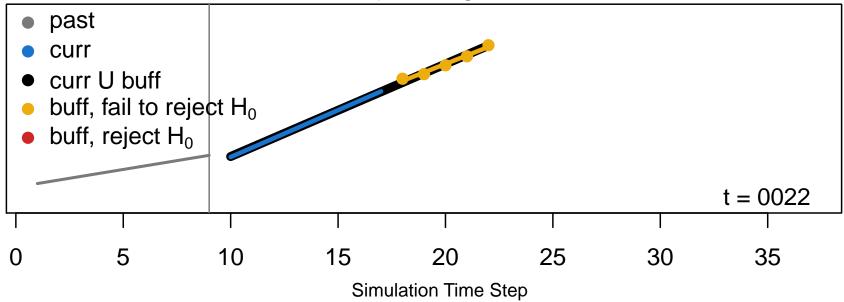




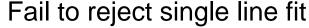


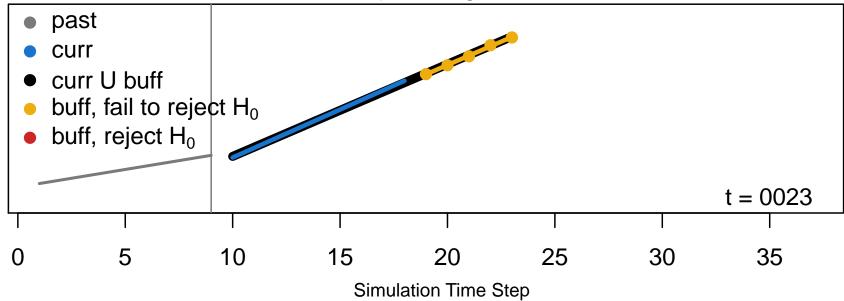




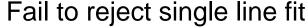


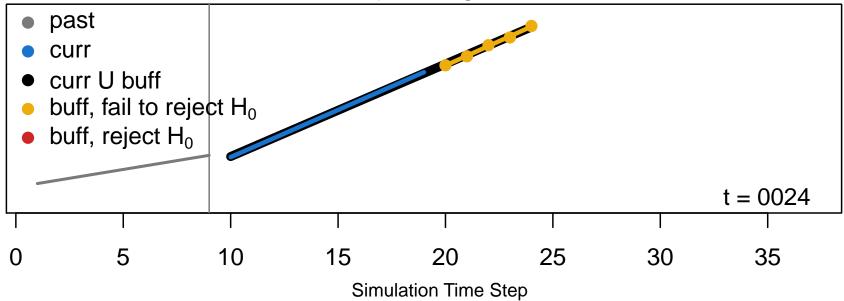




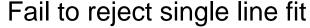


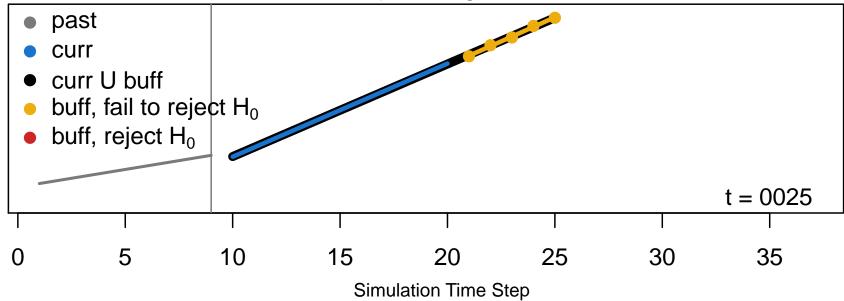




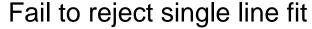


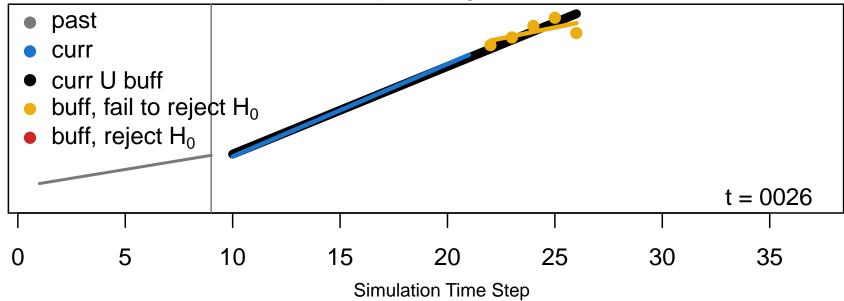




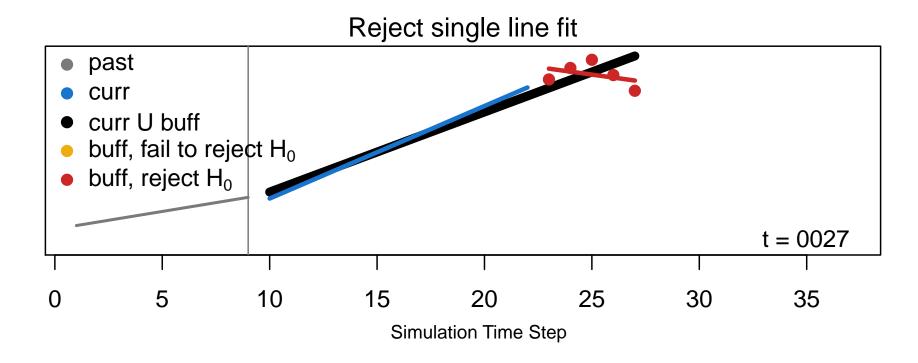




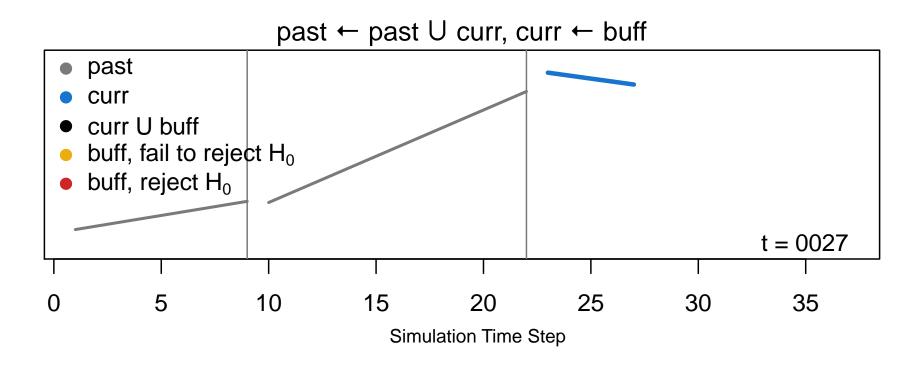




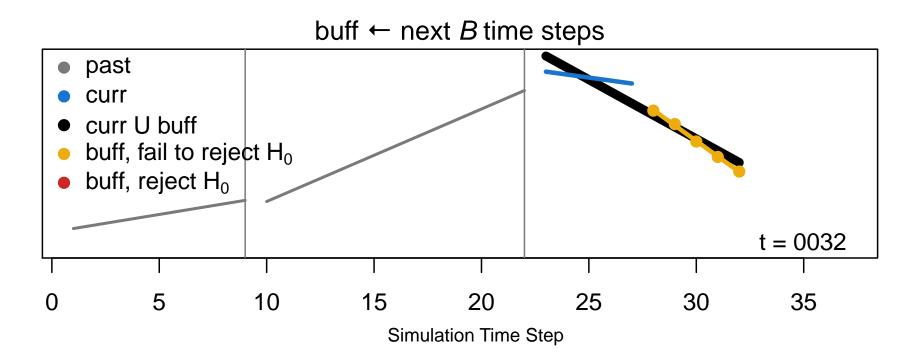






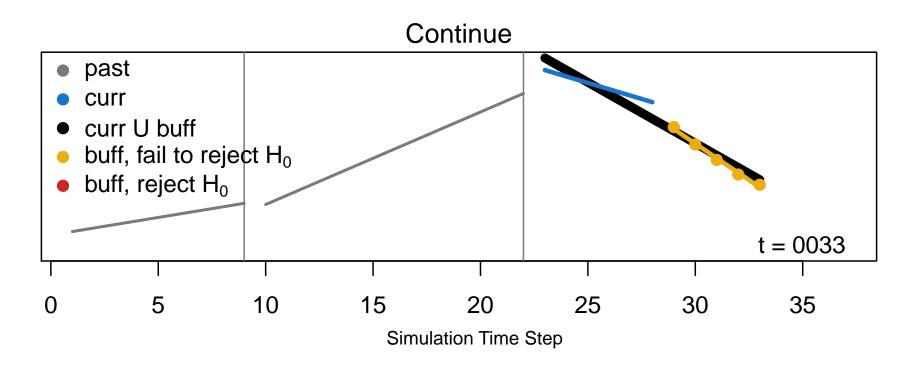








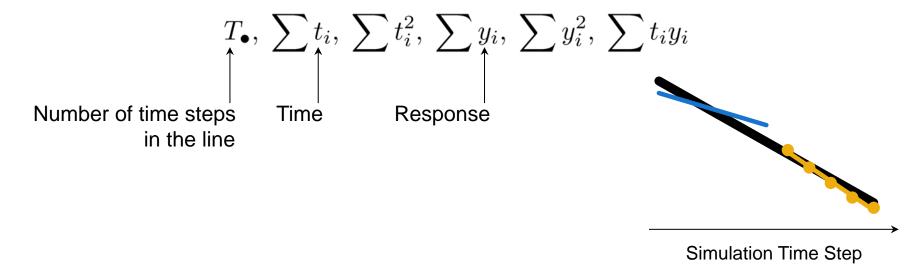








We capture each of the 3 lines with a set of **sufficient statistics**:





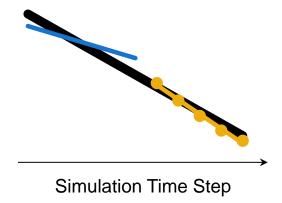


# Our in situ approach: Compare linear fits

We capture each of the 3 lines with a set of **sufficient statistics**:

$$T_{\bullet}, \sum t_i, \sum t_i^2, \sum y_i, \sum y_i^2, \sum t_i y_i$$

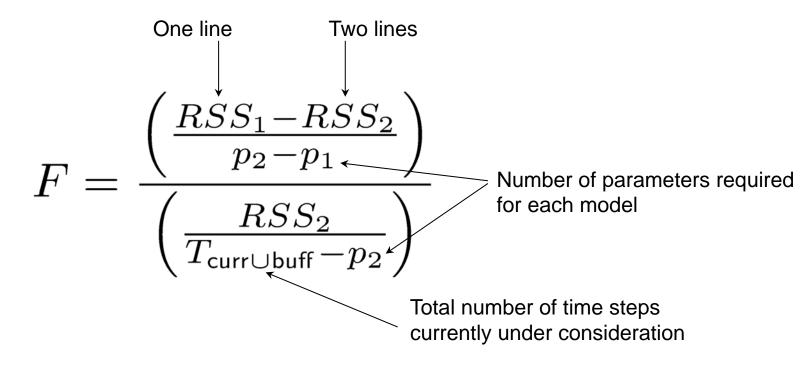
- Update these in constant time, O(1), as the simulation progresses.
- Use to compute the modified F-statistic for our hypothesis test.
- Use to construct a linear approximation of the entire simulation with known error.





### Our modified *F*-statistic

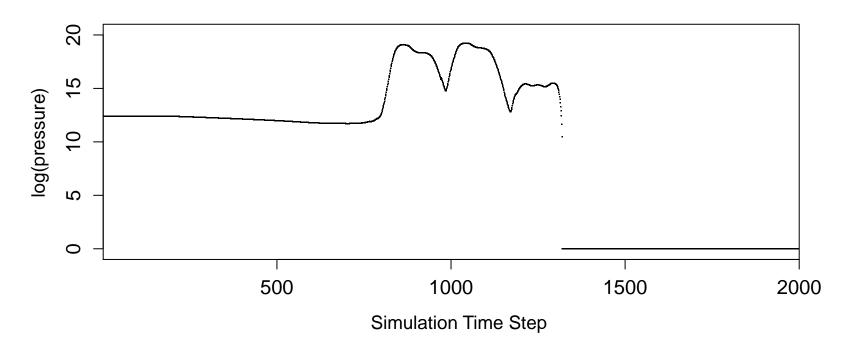
#### Here's the standard formulation:





## Our modified F-statistic

But this can reject  $H_0$  when both curr and buff have extremely low RSS, which is common in these computer simulations.







### Our modified F-statistic

So we add a "nugget"  $\delta^2$ , scaled by  $T_{\text{curr U buff}}$ , to have the effect of adding white noise and encouraging less (or smarter) rejection.

$$F = \frac{\left(\frac{RSS_1 - RSS_2}{p_2 - p_1}\right)}{\left(\frac{RSS_2}{T_{\text{curr} \cup \text{buff}} - p_2}\right)} + T_{\text{curr} \cup \text{buff}} \times \delta^2$$

Now we have 3 "tuning parameters":  $\alpha$ , nugget  $\delta^2$ , and buffer size B. I'll come back to this later. But first: A demo!



## Demo: Is there water on the moon? NASA finds out!

2009 LCROSS Mission: Lunar CRater Observation and Sensing Satellite





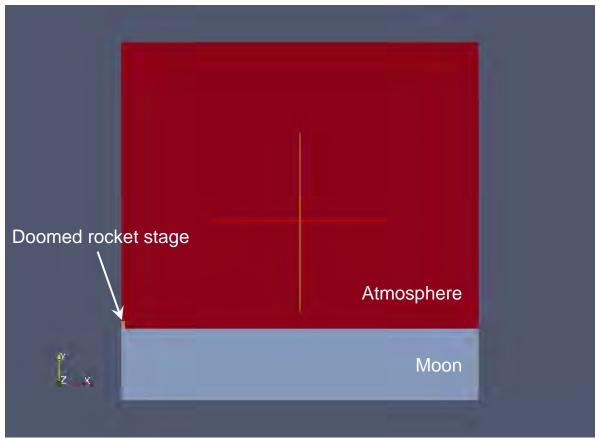






## **But before NASA crashed the Moon...**

Scientists used RAGE simulations to bound the expected results. *Korycansky et al.* 2009





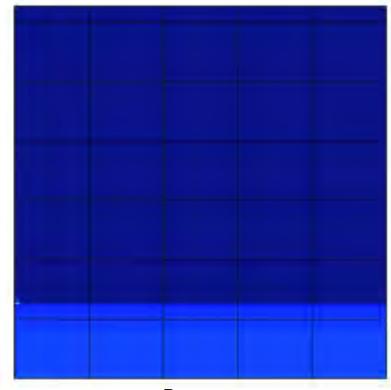


### **But before NASA crashed the Moon...**

Scientists used RAGE simulations to bound the expected results.

Korycansky et al. 2009

- RAGE: A massively parallel Eulerian code used to solve 1D, 2D, or 3D hydrodynamics problems. Gittings et al. 2008
- 2000 time steps, ~10 variables in 2D.
- Not a billion billion calculations per second, but a useful testbed.

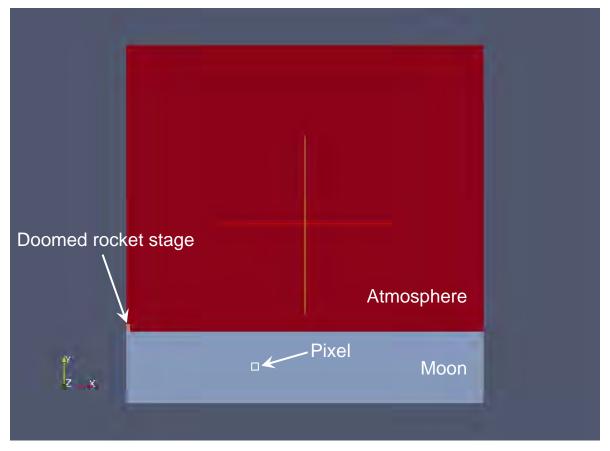


Pressure



## **Demonstration with LCROSS simulation**

First we'll track a pixel of the pressure variable.

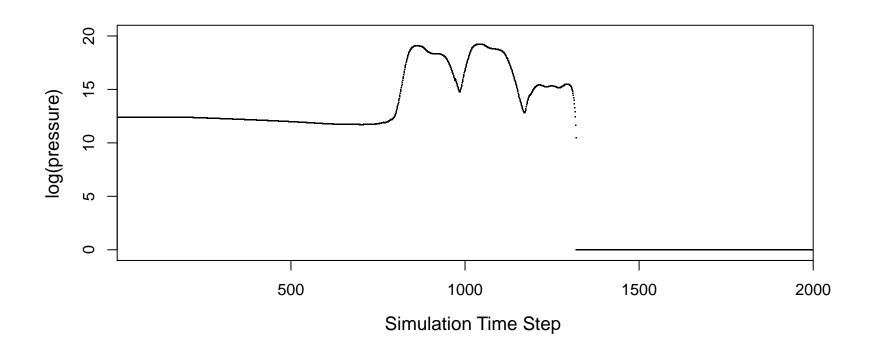






# **Demonstration with LCROSS simulation (single pixel)**

First we'll track a pixel of the pressure variable.

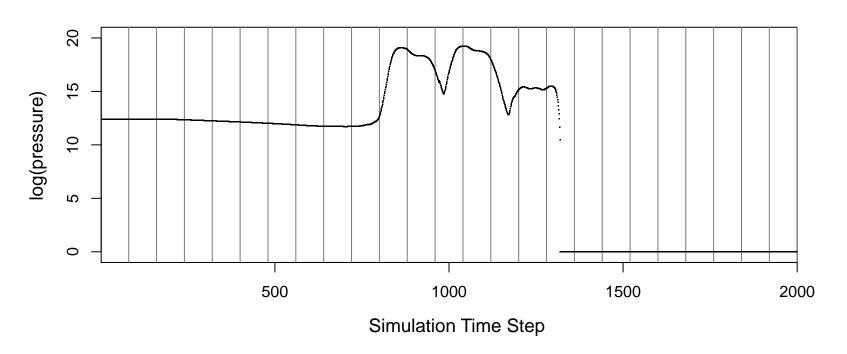






# Demonstration with LCROSS simulation (single pixel)

Standard practice: 25 evenly spaced partitions.





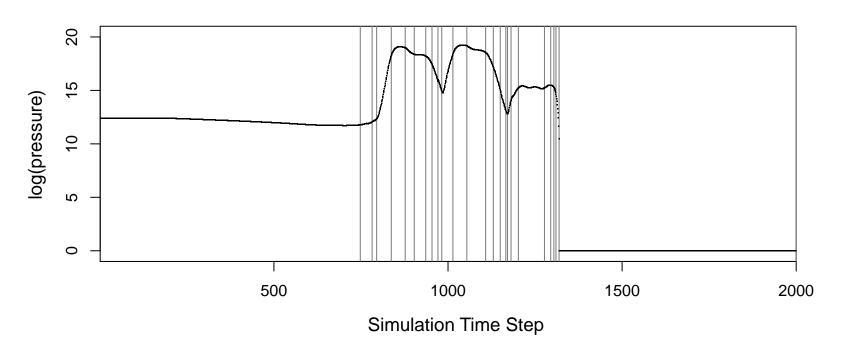
Assuming linear interpolation.

Total RSS: 1140.15



# Demonstration with LCROSS simulation (single pixel)

Our approach: 25 partitions selected with  $\alpha = 0.001$ ,  $\delta^2 = 0.001$ , B = 5.







# But how to choose those tuning parameters?

We've got  $\alpha$ , nugget  $\delta^2$ , and buffer size B.

- B we have little control over.
- We can explore α and δ² in terms of their impact on the number of partitions and the total RSS.

#### Number of partitions

0.1 -	221	210	191	166	145	129	111	84	48	27	13
0.01 -	203	190	173	148	131	115	94	62	40	17	11
0.001 -	173	167	150	130	116	96	73	47	25	16	9
1e-04 -	120	115	105	84	73	62	42	28	18	11	9
α 1e-05 -	60	64	46	38	35	31	29	15	12	11	9
1e-06 -	30	30	27	26	25	23	21	15	11	10	5
1e-07 -	20	18	18	18	17	15	14	11	11	3	5
1e-08 -	11	10	10	10	9	10	10	7	7	3	3
1e-09 -	10	10	10	10	9	9	9	7	7	3	3
1e-10 -	10	10	10	10	9	9	9	4	7	3	3
	0	1e-10 -	1e-09 -	1e-08 -	1e-07 -	1e-06 <sup>-</sup>	1e-05 -	1e-04	0.001	0.01	1.0



 $\delta^2$ 



# But how to choose those tuning parameters?

We've got  $\alpha$ , nugget  $\delta^2$ , and buffer size B.

- B we have little control over.
- We can explore α and δ² in terms of their impact on the number of partitions and the total RSS.

#### Total RSS (rounded)

0.1 -	34	34	34	34	34	0	0	34	1	32	14
0.01 -	0	0	0	0	0	0	1	1	1	8	419
0.001 -	1	1	1	1	1	2	2	1	6	11	534
1e-04 -	1	1	1	1	1	2	4	11	12	282	154
α 1e-05 -	11	11	11	11	11	13	9	44	45	40	154
1e-06 -	9	9	9	9	9	9	9	15	39	307	940
1e-07 <b>-</b>	36	36	36	36	36	36	36	38	38	2709	1027
1e-08 <b>-</b>	1205	1205	1205	1205	1205	1205	1205	1208	1208	2678	2014
1e-09 <b>-</b>	1205	1205	1205	1205	1205	1205	1205	1208	1193	2615	1980
1e-10 -	1205	1205	1205	1205	1205	1205	1205	2078	1194	2584	1980
	0	1e-10 -	1e-09 -	1e-08 -	1e-07 -	1e-06 -	1e-05 -	1e-04		0.01	



 $\delta^2$ 



# Start by understanding the $\delta^2 = 0$ case

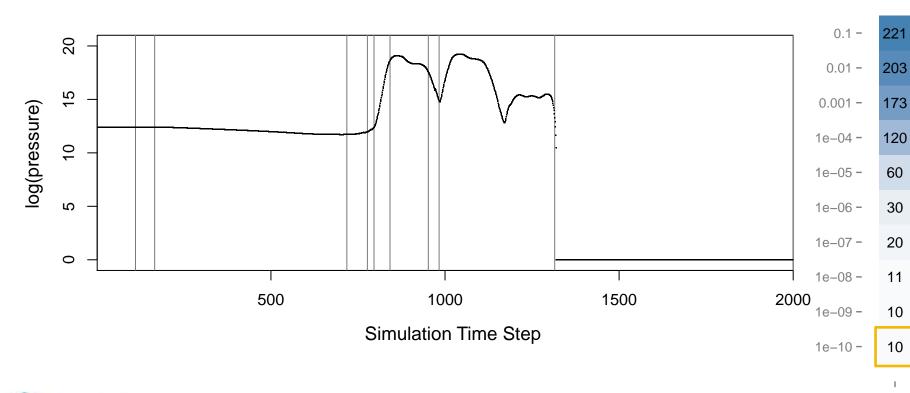
You might think we could just turn the  $\alpha$  knob to reject less often.

#### Number of partitions

							1					
	0.1 -	221	210	191	166	145	129	111	84	48	27	13
α	0.01 -	203	190	173	148	131	115	94	62	40	17	11
	0.001 -	173	167	150	130	116	96	73	47	25	16	9
	1e-04 -	120	115	105	84	73	62	42	28	18	11	9
	1e-05 -	60	64	46	38	35	31	29	15	12	11	9
	1e-06 -	30	30	27	26	25	23	21	15	11	10	5
	1e-07 -	20	18	18	18	17	15	14	11	11	3	5
	1e-08 -	11	10	10	10	9	10	10	7	7	3	3
	1e-09 -	10	10	10	10	9	9	9	7	7	3	3
	1e-10 -	10	10	10	10	9	9	9	4	7	3	3
		ı	1 0	ı	I		I (O	I I	1	_	ı	ı
		0	1e-1	1e-0	1e-0	1e-0	1e-0(	1e-05	1e-0	0.00	0.01	0.1
							$\delta^2$					

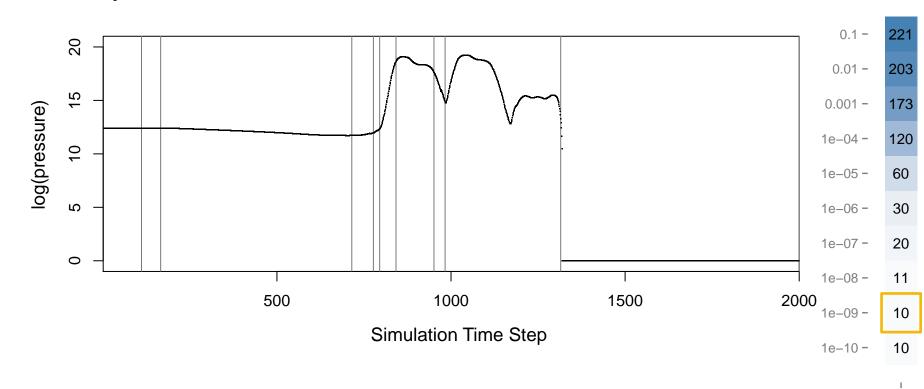


With  $\delta^2 = 0$ , the hypothesis test gets fooled when curr and buff both have extremely low error.

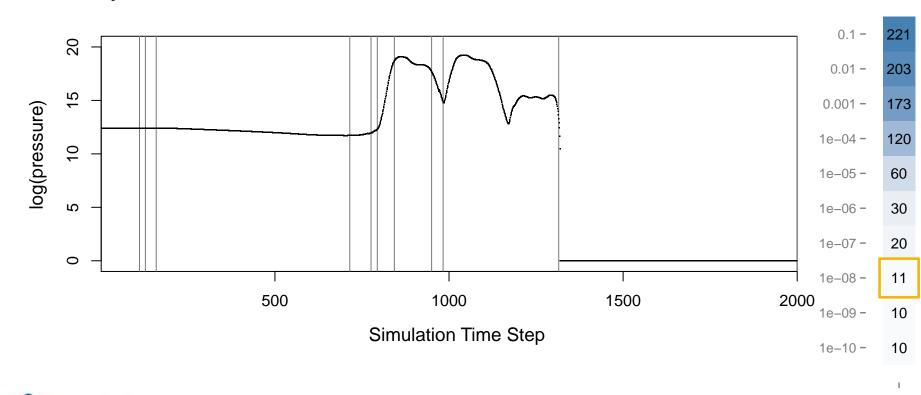




With  $\delta^2 = 0$ , the hypothesis test gets fooled when curr and buff both have extremely low error.

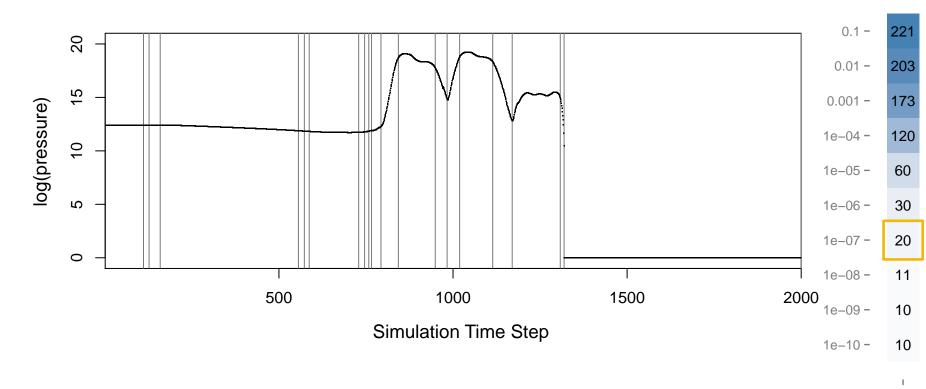




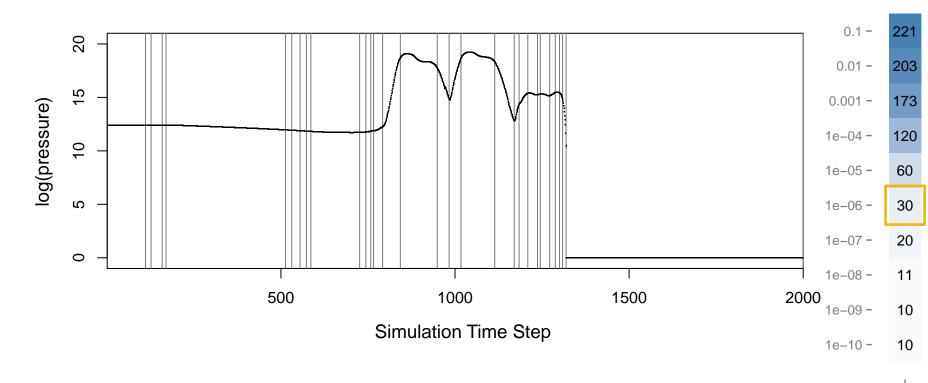




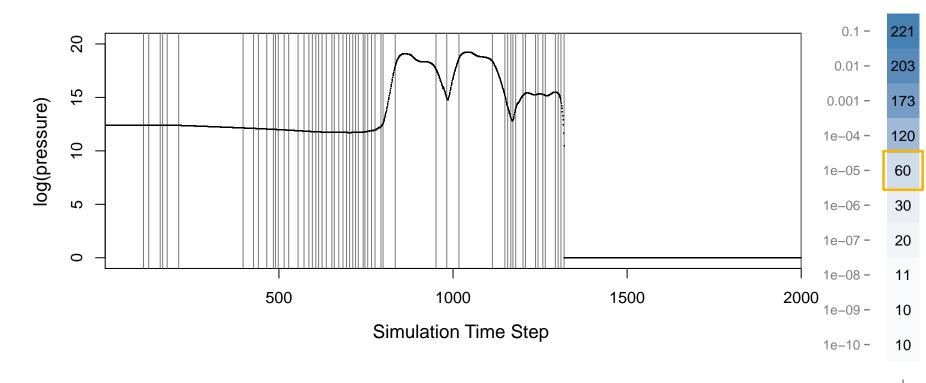
With  $\delta^2 = 0$ , the hypothesis test gets fooled when **curr** and **buff** both have extremely low error.



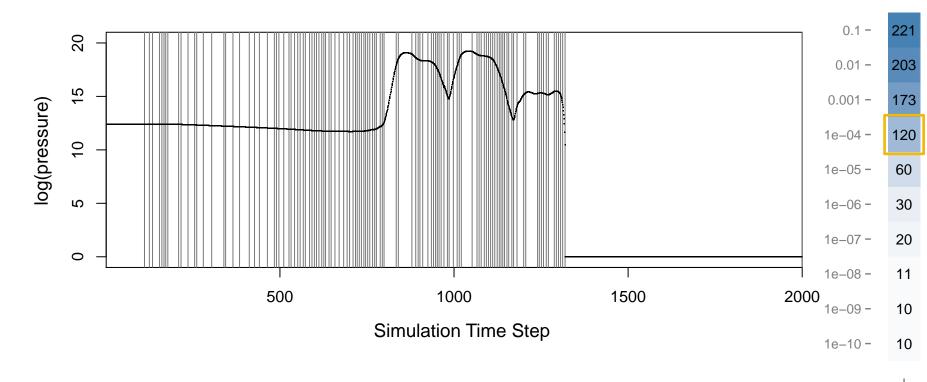




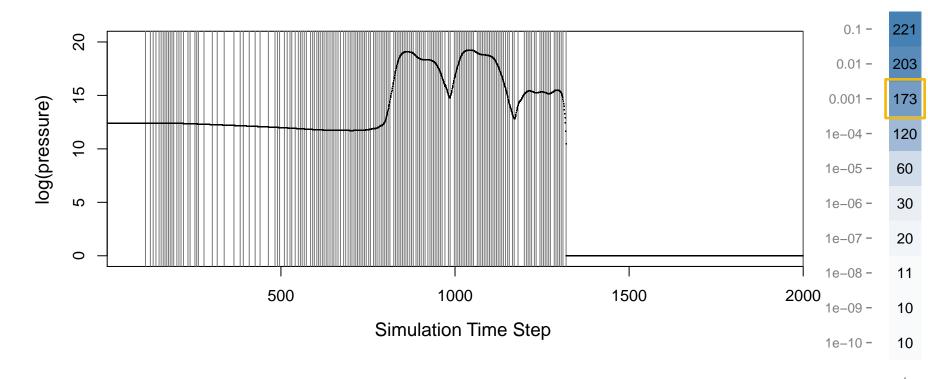




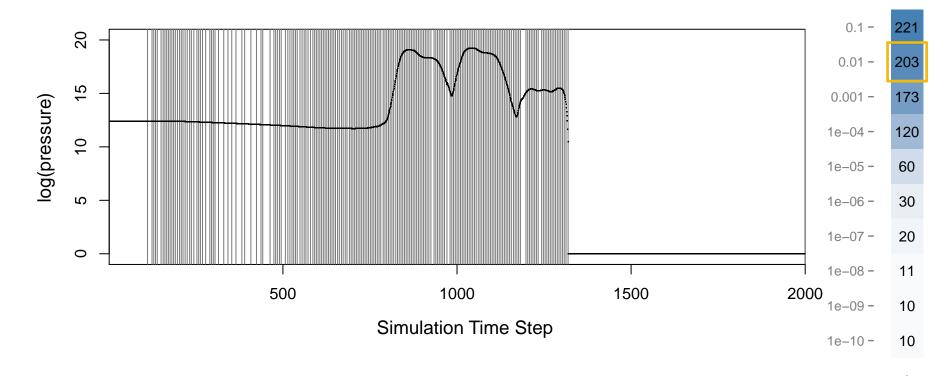




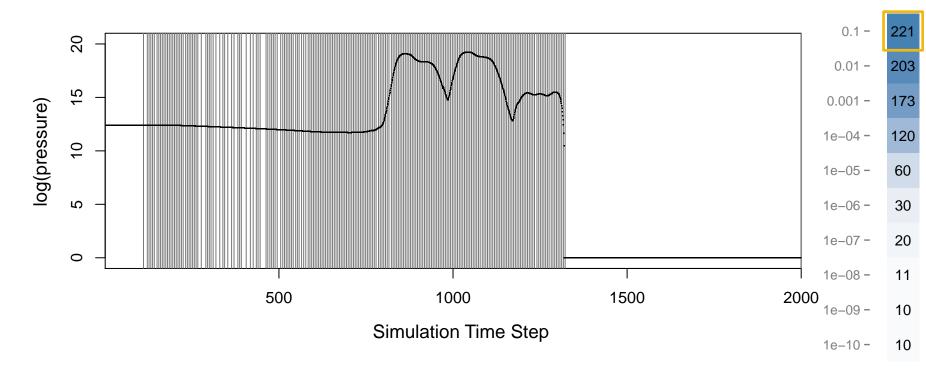










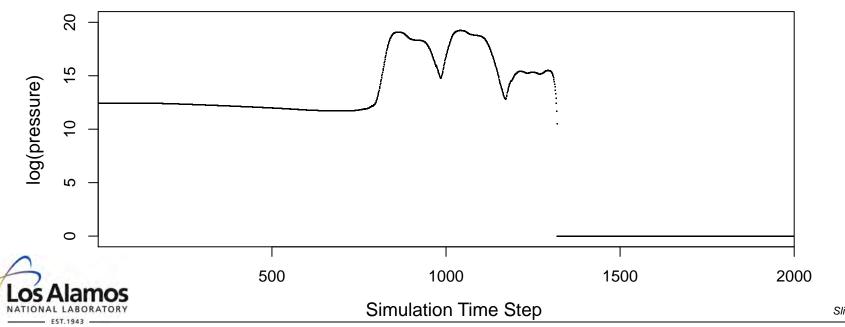




# Q: What's going wrong? A: What isn't going wrong?

These deterministic computer codes violate pretty much every statistical assumption we typically like to make:

- Samples aren't i.i.d. but rather come from a smooth process.
- Error isn't Gaussian.
- Variances of curr and buff aren't necessarily equal.



# Take a look at a positive $\delta^2$ case

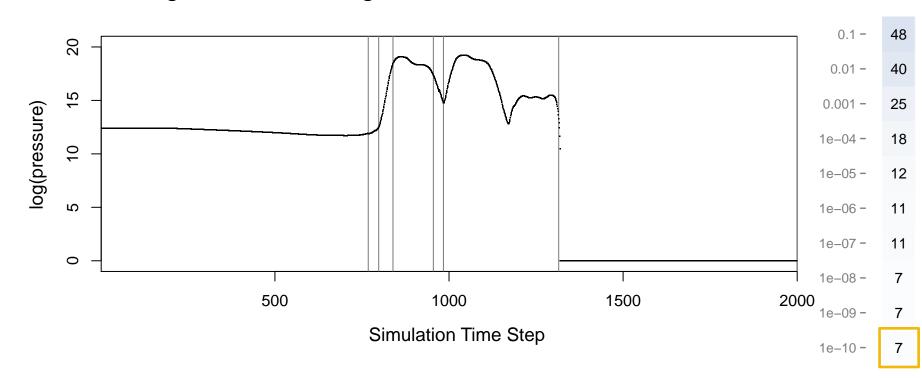
### Number of partitions

							•			_		
	0.1 -	221	210	191	166	145	129	111	84	48	27	13
	0.01 -	203	190	173	148	131	115	94	62	40	17	11
	0.001 -	173	167	150	130	116	96	73	47	25	16	9
	1e-04 -	120	115	105	84	73	62	42	28	18	11	9
α	1e-05 -	60	64	46	38	35	31	29	15	12	11	9
	1e-06 -	30	30	27	26	25	23	21	15	11	10	5
	1e-07 -	20	18	18	18	17	15	14	11	11	3	5
	1e-08 -	11	10	10	10	9	10	10	7	7	3	3
	1e-09 -	10	10	10	10	9	9	9	7	7	3	3
	1e-10 -	10	10	10	10	9	9	9	4	7	3	3
		ı	_ 0	L 60	1 8		_ 90	_ 9	- 4	1	ı_	ı
		0	1e-10	1e-09	1e-08	1e-07	1e-06	1e-05	1e-04	0.001	0.01	0.1
							$\delta^2$					



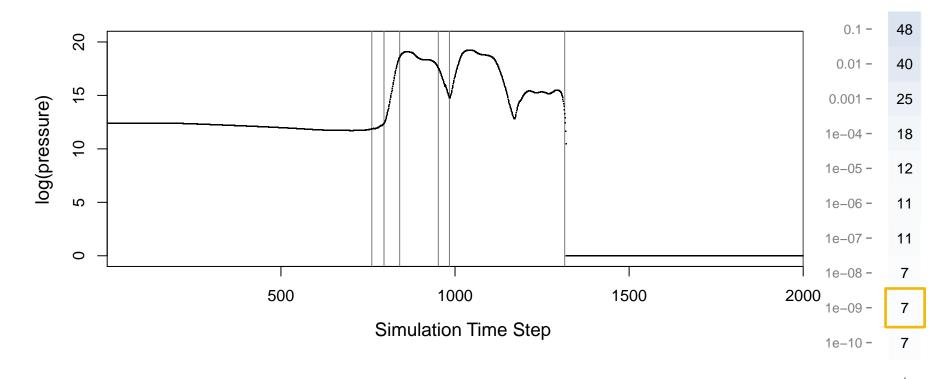


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



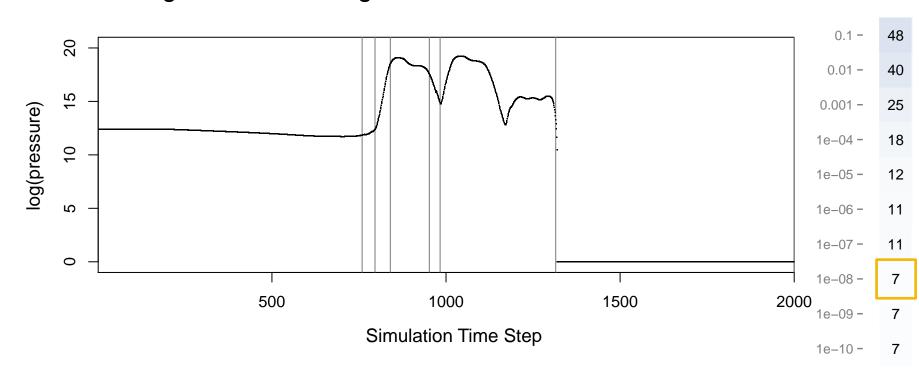


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



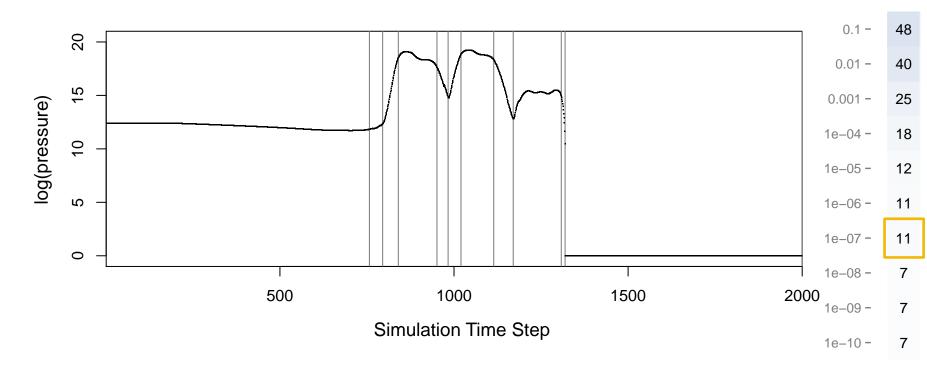


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



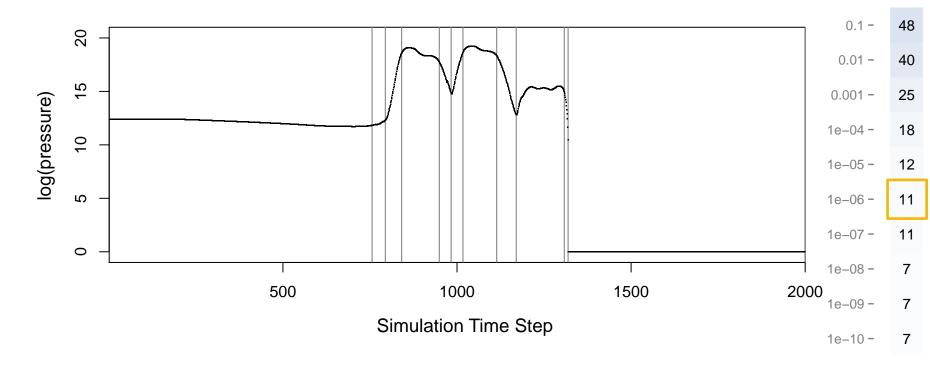


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



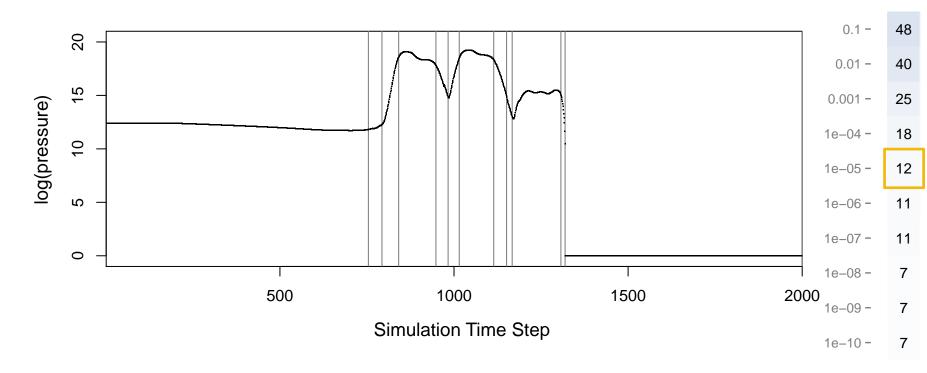


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



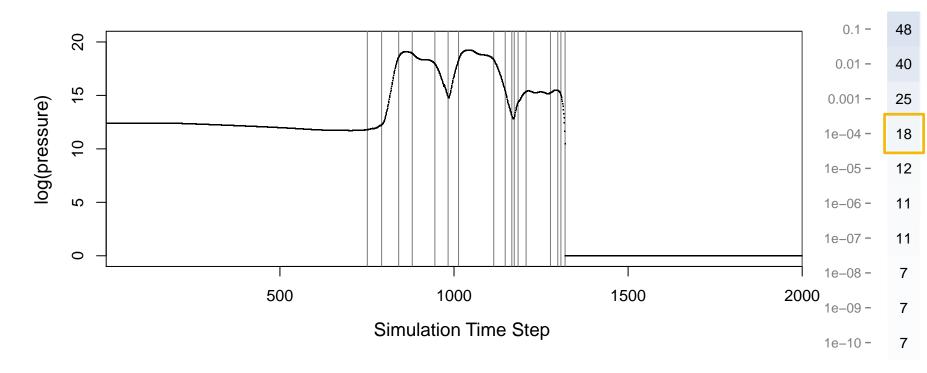


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



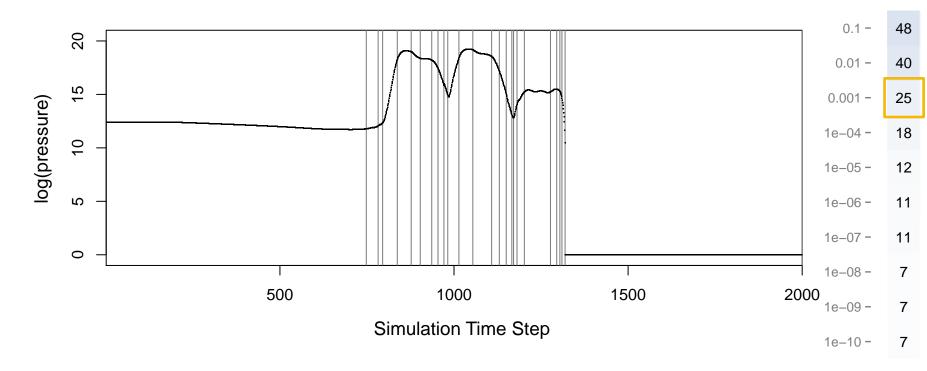


It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.



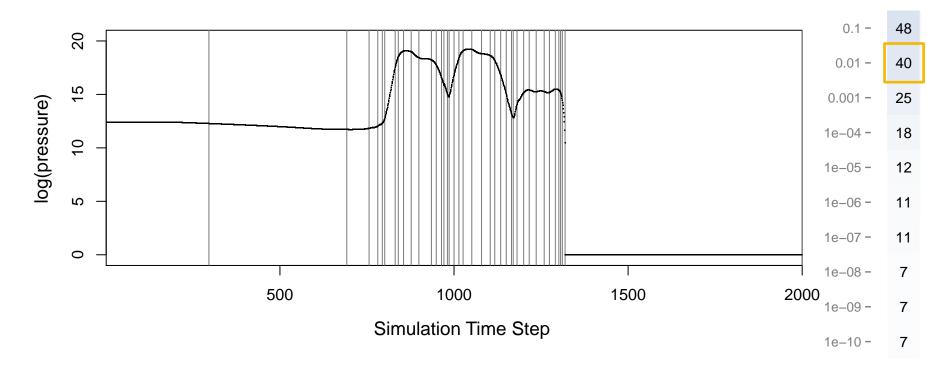


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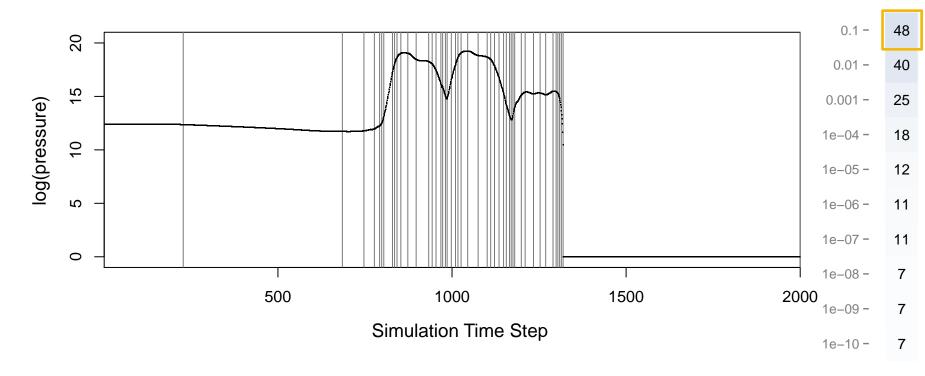
It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.







It's like adding white noise with variance  $\delta^2$ , providing a global effect on the kinds of changes that can be ignored.







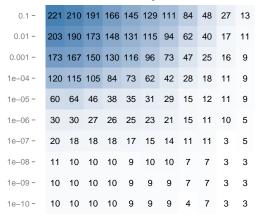
## Ok, but still: How to choose those tuning parameters?

#### Here's what we've learned:

- α governs local choices about partitioning curr and buff. Increasing α fills in areas that already have partitions, making the fit more detailed.
- $\delta^2$  provides a global effect about the kinds of changes to ignore.
- For now we recommend doing a few "scanning" runs of the simulation to build small versions of tables like these.

# Los Alamos NATIONAL LABORATORY

#### Number of partitions



#### **Total RSS**

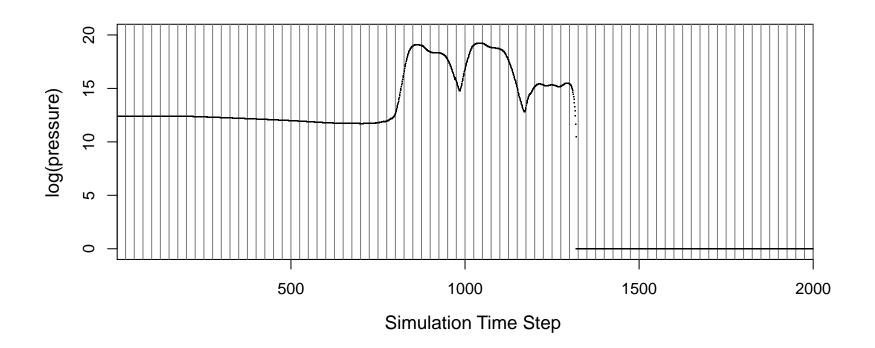


Slide 73



## We argue: It's worth it to do a few scanning runs

Standard practice: 80 partitions, total RSS 280.43.

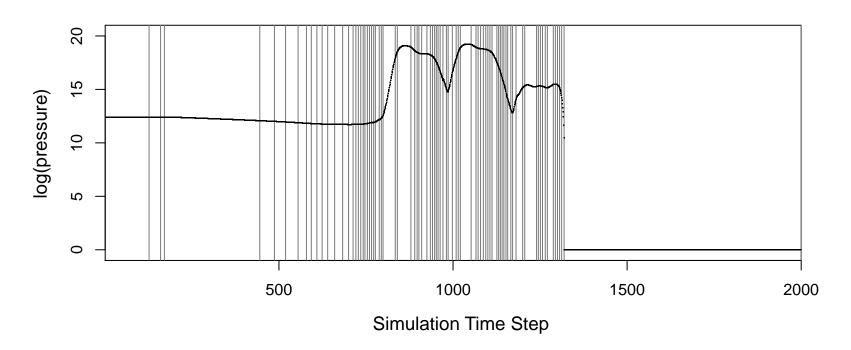






## We argue: It's worth it to do a few extra runs

**Standard practice:** 80 partitions, total RSS 280.43. **Our approach:** 84 partitions, total RSS 1.30.

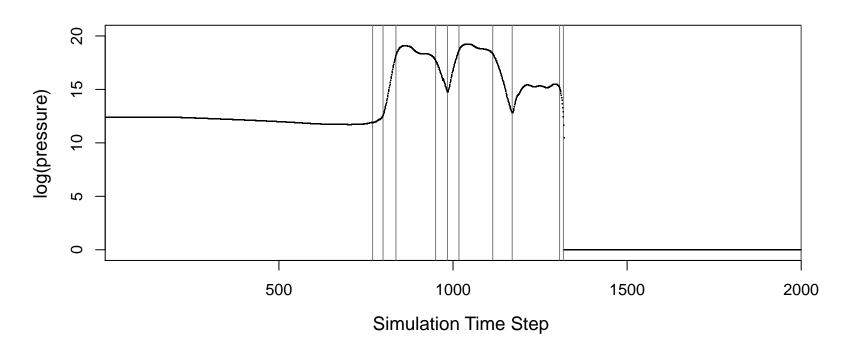




## We argue: It's worth it to do a few extra runs

Standard practice: 80 partitions, total RSS 280.43.

Our approach: 11 partitions, total RSS 281.61.





## Tradeoff between number of partitions and total RSS

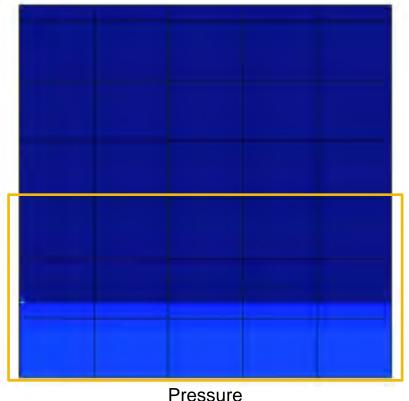
Ultimately would like to find an AIC-like criterion to balance this.

	Number of partitions														Total RSS (rounded)										
	0.1 -	221	210	191	166	145	129	111	84	48	27	13	0.1 -	34	34	34	34	34	0	0	34	1	32	14	
a	0.01 -	203	190	173	148	131	115	94	62	40	17	11	0.01 -	0	0	0	0	0	0	1	1	1	8	419	
	0.001 -	173	167	150	130	116	96	73	47	25	16	9	0.001 -	1	1	1	1	1	2	2	1	6	11	534	
	1e-04 -	120	115	105	84	73	62	42	28	18	11	9	1e-04 -	1	1	1	1	1	2	4	11	12	282	154	
	1e-05 -	60	64	46	38	35	31	29	15	12	11	9	1e-05 -	11	11	11	11	11	13	9	44	45	40	154	
	1e-06 -	30	30	27	26	25	23	21	15	11	10	5	1e-06 -	9	9	9	9	9	9	9	15	39	307	940	
	1e-07 -	20	18	18	18	17	15	14	11	11	3	5	1e-07 -	36	36	36	36	36	36	36	38	38	2709	1027	
	1e-08 -	11	10	10	10	9	10	10	7	7	3	3	1e-08 -	1205	1205	1205	1205	1205	1205	1205	1208	1208	2678	2014	
	1e-09 -	10	10	10	10	9	9	9	7	7	3	3	1e-09 -	1205	1205	1205	1205	1205	1205	1205	1208	1193	2615	1980	
	1e-10 -	10	10	10	10	9	9	9	4	7	3	3	1e-10 -	1205	1205	1205	1205	1205	1205	1205	2078	1194	2584	1980	
		0	1e-10 -	1e-09 -	1e-08 -	1e-07 -	1e-06 -	1e-05 -	1e-04	0.001	0.01	0.1		0	1e-10 -	1e-09 -	16-08	1e-07 -	1e-06 -	1e-05 <sup>-</sup>	1e-04	0.001 -	0.01	0.1	
$oldsymbol{\delta}^2$												$oldsymbol{\delta}^2$						Sli	de 77						

## Incorporating spatial characteristics of the simulation

A simple initial approach with the LCROSS simulation:

- Split the simulation frames into blocks.
- For each block and each time step, compute the mean over pixels.
- Apply the method to the trace of each block mean.

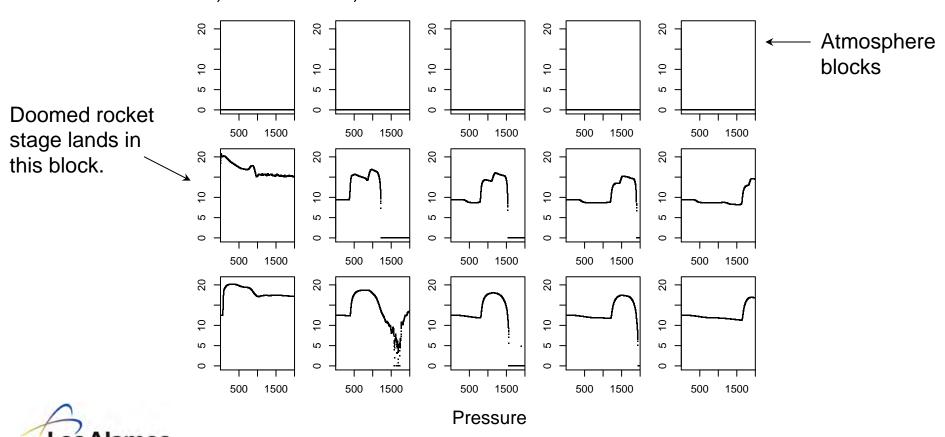






## Incorporating spatial characteristics of the simulation

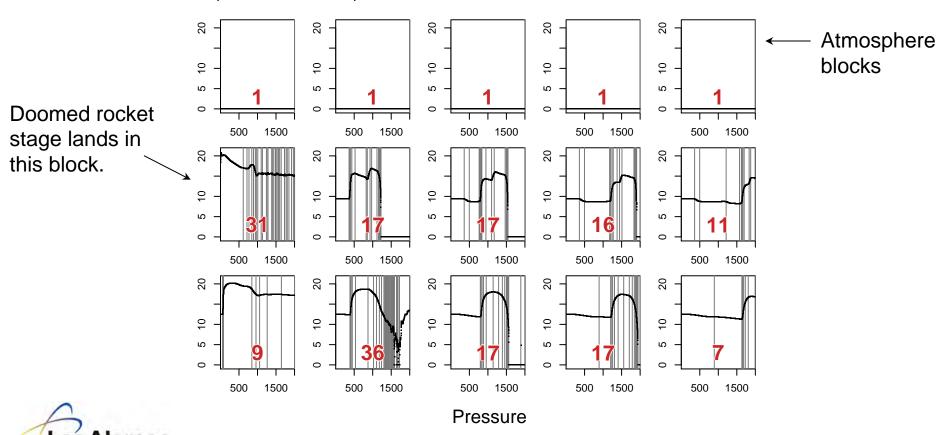
Applying our approach to pixel means for different regions of the simulation with  $\alpha = 0.001$ ,  $\delta^2 = 0.001$ , B = 5.





## Incorporating spatial characteristics of the simulation

Applying our approach to pixel means for different regions of the simulation with  $\alpha = 0.001$ ,  $\delta^2 = 0.001$ , B = 5.

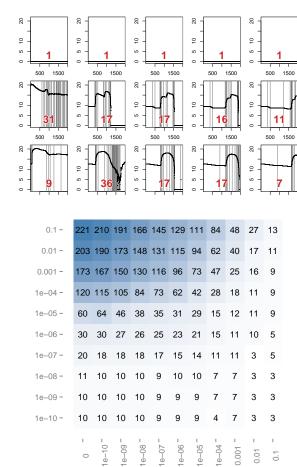




## Lots of potential next directions

#### To name just a few:

- Using our partitioning approach to define spatial regions as the simulation evolves.
- Identifying a mathematical criterion to guide selection of  $\alpha$  and  $\delta^2$ .
- Incorporating other types of fits that can be cheaply computed and updated.
- Handling multivariate trajectories.



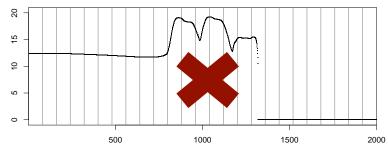


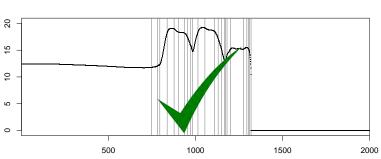
Slide 81

## But for now...

...our in situ approach is cheap to compute and update, and it provides:

- Substantial memory savings over storing the full output of the simulation.
- Improved fidelity to the simulation over selecting evenly spaced partitions.
- Ability to reconstruct a linear approximation of the simulation with known error.









# The end

More details: arxiv.org/abs/1409.0909





# Or: The end!

More details: arxiv.org/abs/1409.0909

Also: Statistics and Beer Day

June 13





### Some math

In a typical simulation setting, a scalar response  $y_i$  will be an unknown deterministic function of time  $t_i$ :

$$y_i = \mathcal{F}(t_i), i = 1, \dots, T$$

where *T* is the total number of time steps in the simulation. Our goal is to approximate this function and locate interesting changes:

$$y_i = f(t_i) + \epsilon_i, i = 1, \dots, T$$

Let  $P_0, P_1, ..., P_m$  be a set of breakpoints of the sequence 1, ..., T, with  $P_0 = 0$  and  $P_m = T$ . The function f can be written as a sum over the partitions defined by the breakpoints:

$$f(t_i) = \sum_{j=1}^{m} (\beta_{j,0} + \beta_{j,1}t_i) I\{P_{j-1} < i \le P_j\}$$

To fit the model, we need to estimate the number of partitions, the breakpoints, and the regression coefficients.



Slide 85

## **Sufficient statistics**

$$\theta = \sum t_i$$

$$\Theta = \sum t_i^2$$

$$\psi = \sum y_i$$

$$\Psi = \sum y_i^2$$

$$\tau = \sum t_i y_i$$

$$T_{\bullet}$$

Compute the residual sum of squares (RSS) and the slope and intercept:

$$RSS = \Psi - \frac{1}{T_{\bullet}} \psi^2 - \frac{(\tau - \theta \psi/T_{\bullet})^2}{\Theta - \theta^2/T_{\bullet}}$$

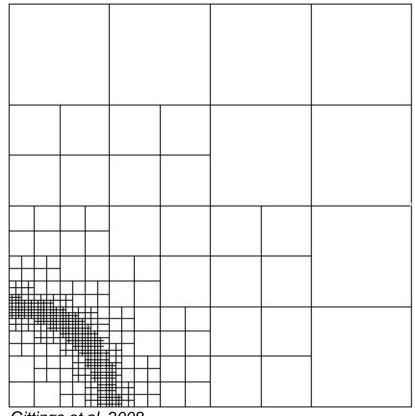
$$\hat{\beta}_0 = \frac{1}{T_{\bullet}} (\psi - \hat{\beta}_1 \theta)$$

$$\hat{\beta}_1 = \frac{\tau - \theta \psi/T_{\bullet}}{\Theta - \theta^2/T_{\bullet}}$$



## RAGE uses adaptive mesh refinement (AMR)

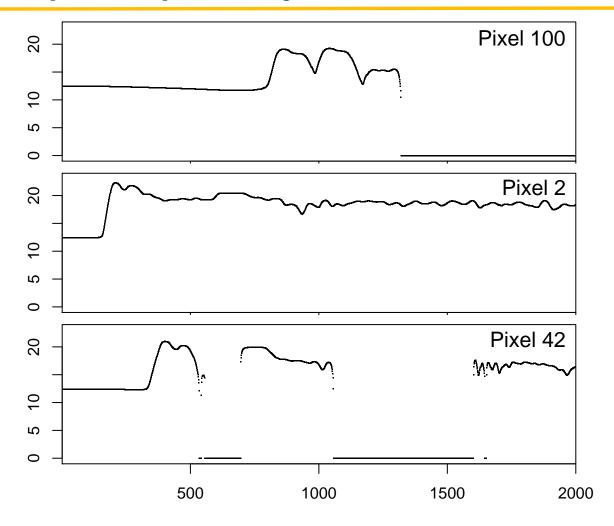
- Considers spatial variation in each variable to choose cell size.
- Makes decisions to split / merge cells at each time step.
- Constrains splits and merges so adjacent cells are within 1 level of each other.



Gittings et al. 2008



## Other examples of pixel trajectories







## We describe capturing the lines with sufficient statistics

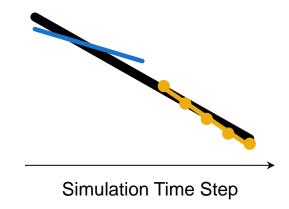
But in practice, these sums can get too large to be computationally stable.

$$T_{\bullet}, \sum t_i, \sum t_i^2, \sum y_i, \sum y_i^2, \sum t_i y_i$$

An alternative: incremental QR decomposition:

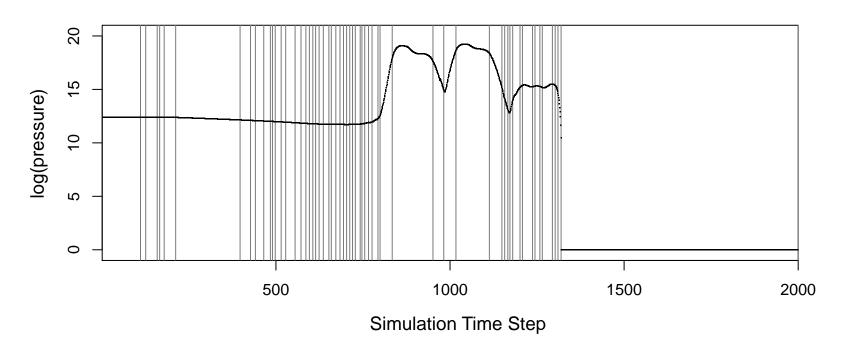
Miller (1992). Algorithm AS 274: Least Squares Routines to Supplement Those of Gentleman, *Journal of the Royal Statistical Society. Series C (Applied Statistics)*, Vol. 41, No. 2, pp. 458-478

This is implemented in the R package biglm.





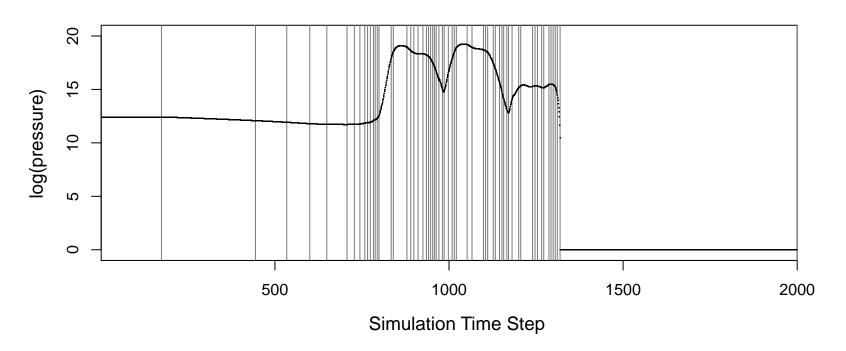
60 partitions via  $\alpha = 1 \times 10^{-5}$ ,  $\delta^2 = 0$ .







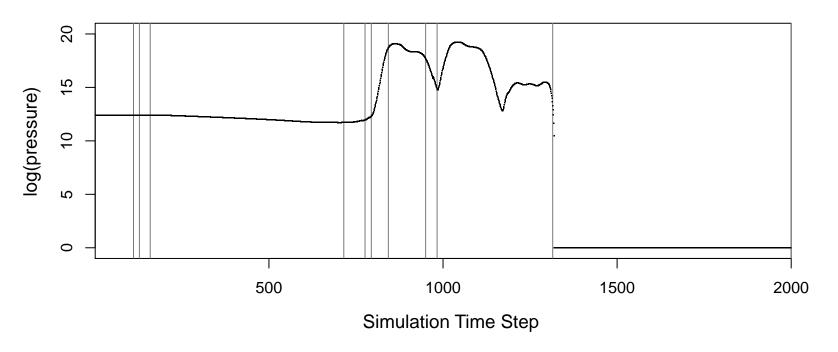
62 partitions via  $\alpha = 1 \times 10^{-4}$ ,  $\delta^2 = 1 \times 10^{-6}$ .







11 partitions via  $\alpha = 1 \times 10^{-8}$ ,  $\delta^2 = 0$ .









11 partitions via  $\alpha = 1 \times 10^{-7}$ ,  $\delta^2 = 1 \times 10^{-4}$ .

